## Homework 8 (Due Thursday, February 23rd)

1. The canonical partition function for a *d*-dimensional phonon gas is

$$\ln Q = -\mathcal{C} \int_0^{\omega_m} \ln(1 - e^{-\beta\hbar\omega}) \,\omega^{d-1} \,d\omega$$

where C and  $\omega_m$  are constants. Show that in the low temperature limit the heat capacity  $C_v$  goes as  $T^d$  while in the high temperature limit it goes to a constant independent of temperature. Note that while most solids vibrate in three dimensions, the lattices in selenium and graphite roughly behave as one-dimensional and two-dimensional phonon gases, respectively.

 $\mathbf{2}$ . (a) The canonical partition function for a photon gas is

$$Q = \exp\left[\frac{4\sigma}{3ck}VT^3\right]$$

where  $\sigma$  is Stefan's constant. Find the internal energy U(T, V) and entropy S(T, V).

(b) The energy density of a photon gas is

$$u(\omega) = \frac{a\omega^3}{\exp(\beta\hbar\omega) - 1}$$

where a is a constant. Accurately plot this function to show that it has a peak at  $\hbar\omega \approx 2.82 \, kT$ .

(c) In 1965, Penzias and Wilson discovered a nearly isotropic cosmic microwave background. The spectrum is approximately that of a blackbody at about 2.8 Kelvins with a maximum energy density at a frequency of about  $\nu_{\text{max}} = 160$  GHz. In a simple view, this radiation comes from the adiabatic expansion of a much hotter photon cloud that was produced in the Big Bang, about 13 billion years ago. What was the peak frequency,  $\nu_{\text{max}}$ , at the time that life was forming on Earth, about 4 billion years ago? Assume that the radial expansion rate of the universe is constant; note that  $\nu = \omega/(2\pi)$ .